AKARI Far-Infrared View of Nearby Galaxies

H. Kaneda¹, T. Suzuki², T. Onaka³, I. Takase³, M. Yamagishi¹, D. Ishihara¹ and I. Sakon³

¹Graduate School of Science, Nagoya University, Chikusa-ku, Nagoya 464-8602, Japan email: kaneda@u.phys.nagoya-u.ac.jp

²National Astronomical Observatory of Japan, Mitaka, Tokyo 181-8588, Japan ³Graduate School of Science, The University of Tokyo, Bunkyo-ku, Tokyo 113-0033, Japan

Abstract. We have observed 57 nearby galaxies in the far-infrared with the Far-Infrared Surveyor on AKARI to study the properties of dust in various environments.

Keywords. ISM: dust, extinction — infrared: galaxies — galaxies: ISM

AKARI far-infrared (IR) observations of 57 nearby galaxies have been carried out in part of the AKARI mission program "ISM in our Galaxy and Nearby Galaxies" (Kaneda et al. 2009a). The Far-IR Surveyor (FIS) on AKARI has 4 photometric bands at the wavelengths of 65, 90, 140, and 160 μ m, which are of great use to accurately determine spatial variations in the properties of dust. For face-on spiral galaxies such as M 101 and M 81, we spectrally decompose dust emission into warm and cool components and spatially resolve each component, whereby we obtain physical insight into relationship between star formation rates and ISM densities (e.g. Suzuki *et al.* 2007). Large dynamic ranges of signal detection of the FIS provided by a special read-out mode is another advantage; we clearly confirm the presence of far-IR dust in the halo of the edge-on starburst galaxy NGC 253 without saturation problems at its very bright nucleus (Kaneda et al. 2009b). For NGC 1316, we find that far-IR emission is extending along the mid-IR jet-like structures, getting softer away from the galactic nucleus (Fig.1). Spatial information on vertical structures as well as face-on ones of galaxies is important in view of material circulation in a galaxy. Our AKARI data will be complementary to both existing and upcoming far-IR data of nearby galaxies such as those from Spitzer and Herschel.

AKARI is a JAXA project with the participation of ESA.



Figure 1. AKARI far-IR 4-band contour maps of NGC 1316, overlaid on the AKARI 11 μ m image. Contour levels are logarithmically drawn from 10 to 80 % of the peak brightness.

References

Kaneda, H., Koo, B.-C., Onaka, T., & Takahashi, H. 2009a, *Adv.Sp.Res.* 44, 1038 Kaneda, H., Yamagishi, M., Suzuki, T., & Onaka, T. 2009b, *ApJ* 698, L125 Suzuki, T. *et al.* 2007, *PASJ* 59, S473

414