Diamagnetic drift instabilities in collisional non-uniform quantum dusty magnetoplasmas

M. JAMIL¹, CH. UZMA¹, K. ZUBIA¹, I. ZEBA¹, H. M. RAFIQUE² and M. SALIMULLAH³

¹Department of Physics, GC University, Lahore-54000, Pakistan (jamil.gcu@gmail.com) ²Department of Physics, University of the Punjab, Lahore-54000, Pakistan

³Department of Physics, Jahangirnagar University, Savar, Dhaka-1342, Bangladesh

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Abstract. Dust-acoustic and dust-lower-hybrid diamagnetic drift wave instabilities have been examined in a collisional non-uniform quantum dusty magnetoplasma. The dust-acoustic drift instability arises through the Fermi degenerate pressure of electrons for a high-density plasma, while for a relatively low-density collisional quantum plasma and short wavelength consideration, the instability is dominated by the Bohm potential effect exciting a new quantum dust-acoustic wave. In the long-range wavelength limit, dust-lower-hybrid waves are found to be unstable because of the diamagnetic drift of magnetized ions. Various possible instability conditions are found for diamagnetic drift instability.

During the last two decades, low-frequency waves and instabilities in dusty plasmas with or without the presence of external magnetic fields have occupied the major areas of dusty plasma physics. This is because all properties of dusty plasmas in general are related with waves and instabilities. With the advent of dustacoustic (DA) and dust-lower-hybrid (DLH) waves (Rao et al. 1990; Salimullah et al. 1992; Shukla 1992; Shukla and Silin 1992; Salimullah 1996; Jamil et al. 2010) involving dust dynamics, a considerable amount of progress has been achieved in the context of astrophysical and laboratory dusty plasma systems. Earlier, cross-field instabilities were investigated in the presence of external electric and magnetic fields in possible laboratory dusty plasma experiments (Shukla et al. 2002; Rastunkov and Krainov 2004; Rosenberg and Shukla 2004; Norreys et al. 2009; Sabeen et al. 2010). In recent years, there has been a growing interest in quantum dusty plasmas because of their importance in micro-electronics and electronic devices with nano-electronic components, dense astrophysical objects, and in laser-produced plasmas (Kremp et al. 1999; Andreev 2000; Jung 2001; Opher et al. 2001; Chabrier et al. 2002; Haas et al. 2003; Bingham et al. 2004; Marklund and Shukla 2006; Shukla et al. 2006a; Brodin and Marklund 2007; Marklund and Brodin 2007; Brodin et al. 2008a; Ren et al. 2009; Salimullah et al. 2009a; Hussain et al. 2010). When a plasma is cooled to an extremely low temperature, the de Broglie wavelengths of plasma particles could be at least comparable to the scale lengths, such as the Debye length or the Larmor radius etc., in the system. In such plasmas, the ultra-cold dense plasma would behave as a Fermi gas, and quantum mechanical effects might play a vital role in the behavior of charge carriers of these plasmas under extreme conditions. There are different models for quantum plasmas. Here we are using the Schrodinger–Poisson model, which is a simplified model and capture the main features of quantum plasmas. It is more appropriate for the analytical and numerical treatment of problems. It makes direct use of global properties of plasma quantities, such as density and average velocity, and therefore may ignore the spin effects (Haas 2005; Haas et al. 2000; Asenjo et al. 2011).

Spin effects of quantized charges are important for linear modes in the presence of a strong magnetic field, and for the systems whose confined geometry is frequently comparable with the thermal de Broglie wavelengths like micro electronics, nano-wires, quantum dots, etc. In large dimensional systems, because of random orientation of spins, high-density collisional plasmas, and the systems where the electrostatic forces are strong as compared with spin forces, the macroscopic spin population may not sustain and therefore become negligible (Garcia et al. 2005; Shukla 2006; Shukla and Eliasson 2006; Brodin et al. 2008b).

Collision between the quantum mechanical identical particles is considerably different from those in classical plasmas because of their collision cross sections. An increase in plasma oscillation frequency in quantum plasmas suppresses the collision probability, e.g., electron–electron collision is negligibly small [Na and Jung 2008].

In this paper, we study the diamagnetic drift instabilities of DA and DLH waves in an inhomogeneous and collisional quantum dusty magnetoplasma by including the collision frequency of ions and dust particles. We consider an infinitely extended inhomogeneous highdensity dusty magnetoplasma containing electrons, ions, and charged dust grains in the presence of a homogeneous static ambient magnetic field $\mathbf{B}_0 \parallel \hat{\mathbf{z}}$. At equilibrium, we assume that the charge quasi-neutrality condition is satisfied, i.e., $n_{i0} + (q_d/e)n_{d0} = n_{e0}$, where $n_{j0}(x)$ is the equilibrium number density of the *j*th species (*j* = electrons, ions, or dust), q_d is the average charge on a dust grain, and *e* is the electronic charge. We analyze the stability properties of the system against electrostatic perturbations, including the effects of heavier species.

The governing equations in the quantum hydrodynamic (QHD) model (Gardner 1994, Gasser et al. 2000, Manfredi and Haas 2001; Haas 2005; Manfredi 2005; Ali and Shukla 2006; Shukla et al. 2006; Ren et al. 2008, Saleem et al. 2008) for electrons, ions, and charged dust grains (j = e, i, d) in the presence of the ambient magnetic field **B**₀ are

$$m_{j} n_{j0} \frac{\partial}{\partial t} \mathbf{v}_{j} = -n_{j0} q_{j} \nabla \phi + n_{j0} \frac{q_{j}}{c} \mathbf{v}_{j} \times \mathbf{B}_{0} - \nabla p_{Fj1} + \frac{\hbar^{2}}{4m_{j}} \nabla (\nabla^{2} n_{j1}).$$
(1)

The corresponding components of velocities for jth species can be written from (1) as

$$\mathbf{v}_{j}^{L} = \frac{q_{j}\Phi}{m_{j}\omega} \left[\frac{i\omega\omega_{cj}}{\omega^{2} - \omega_{cj}^{2}} k_{y}\hat{x} + \frac{\omega^{2}}{\omega^{2} - \omega_{cj}^{2}} k_{y}\hat{y} + k_{z}\hat{z} \right], \quad (2)$$

where

$$\Phi = \phi + \frac{m_j}{q_j} V_{Fj}^{\prime 2} \frac{n_{j1}}{n_{j0}},$$

and

$$V'_{Fj} = V_{Fj}(1+\gamma_j)^{1/2}, \quad \gamma_j = \hbar^2 k^2 / 8m_e K_B T_{Fj},$$

where $n_j = n_{j0} + n_{j1}$, \hbar is the Planck's constant divided by 2π , $\phi(\mathbf{r}, t)$ is the electrostatic potential in the quantum magnetoplasma, and q_j , m_j , n_j , and c are the charge, mass, total equilibrium number density with equilibrium value n_{j0} of the *j*th species, and the velocity of light in vacuum, respectively. Here we take into account the quantum effects of all the species in general when they are considered extremely cold. In (1), we assume that the plasma particles in a zero-temperature Fermi gas satisfy the pressure law (Manfredi and Haas 2001; Manfredi 2005), $p_{Fj} = m_j V_{Fj}^2 n_j^3 / 3n_{j0}^2$, where $V_{Fj} = (2K_B T_{Fj}/m_j)^{1/2}$ is the Fermi speed; K_B , T_{Fj} , and n_j are the Boltzmann constant, the Fermi temperature, and the total number density with its equilibrium value n_{j0} , respectively.

It may be mentioned that in a three-dimensional quantum plasma, the total pressure should be proportional to $n^{(N+2)/N}$, where N = 3. However, according to Manfredi and Haas (2001) this choice is not good, as the results of QHD model differ from the Wigner-

Poisson model. Therefore, the total pressure of the threedimensional quantum plasma is approximately described by the one-dimensional Fermi pressure. Similar pressure law was employed by Ali and Shukla (2006) for threedimensional quantum plasmas.

The equation of continuity for inhomogeneity along x-direction is

$$\frac{\partial n_j}{\partial t} + \nabla . (n_j \mathbf{v}_j) = 0.$$
(3)

After linearization of continuity equation, we have

$$\frac{\partial n_{j1}}{\partial t} + n_{j0} (\nabla \cdot \mathbf{v}_{jy} + \nabla \cdot \mathbf{v}_{jz}) + \mathbf{v}_{jx} \cdot \partial n_{j0} / \partial x = 0.$$
(4)

The Fourier transformation of (4) gives the following equation:

$$n_{j1}^{L} = \frac{n_{j0}}{\omega} k_{z} v_{jz}^{L} + \frac{n_{j0}}{\omega} k_{y} v_{jy}^{L} + \frac{n_{j0}}{i \, \omega} v_{jx}^{L}.$$
 (5)

Further, the Poisson's equation satisfying the electrostatic potential ϕ of electrostatic perturbation is

$$\nabla^2 \phi = 4\pi e \left(n_{e1} - n_{i1} - \frac{q_d}{e} n_{d1} \right).$$
 (6)

In the presence of density inhomogeneities in xdirection and the ambient magnetic field, $\mathbf{B}_0 = \hat{\mathbf{z}} B_0$, we assume the presence of drift waves propagating in the yz-plane, proportional to $\exp[-i(\omega t - k_y y - k_z z)]$, where $k^2 = k_y^2 + k_z^2$ and $k_y^2 \gg k_z^2$. Here ω and \mathbf{k} are angular frequency and wavenumber vector, respectively.

We would like to mention that for weak inhomogeneity approximation (Ren et al. 2008) with $k_x \ll k$ and $k \ge 1/L_j$, we can retain the dominant linear term in the Bohm potential term of the equation of motion. Similar consideration has been employed earlier in the literature (Shukla et al. 2006b) for drift wave investigations in the non-uniform quantum magnetoplasmas. However, the non-uniformity of equilibrium density is taken through the continuity equation to study drift waves in nonuniform quantum plasmas.

Using velocity components from (2) into (5), we have

$$n_{j1}^{L} = \frac{n_{j0}q_{j}}{\omega} \left[\frac{k_{z}^{2}}{\omega^{2}} + \frac{k_{y}^{2}}{\omega^{2} - \omega_{cj}^{2}} \left(1 - \frac{\omega_{cj}}{k_{y}L_{j}\omega} \right) \right] \Phi.$$
(7)

After putting Φ

$$n_{j1}^{L} = \frac{\frac{n_{j0}q_{j}}{\omega} \left[\frac{k_{z}^{2}}{\omega^{2}} + \frac{k_{y}^{2}}{\omega^{2} - \omega_{cj}^{2}} \left(1 - \frac{\omega_{cj}}{k_{y}L_{j}\omega}\right)\right]\phi}{1 - V_{Fj}^{\prime 2} \left[\frac{k_{z}^{2}}{\omega^{2}} + \frac{k_{y}^{2}}{\omega^{2} - \omega_{cj}^{2}} \left(1 - \frac{\omega_{cj}}{k_{y}L_{j}\omega}\right)\right]}.$$
(8)

We know,

$$a_{j1}^{L} = -\frac{1}{4\pi q_{j}} \chi_{j} k^{2} \phi.$$
(9)

From (8) and (9), we obtain the dielectric susceptibility for *j*th species, where j = e, i, d as (Baines et al. 1965; Salimullah et al. 2009b)

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$$\chi_{j} = -\frac{\omega_{pj}^{2} \left[\frac{k_{z}^{2}}{\omega^{2}} + \frac{k_{y}^{2}}{\omega^{2} - \omega_{cj}^{2}} \left(1 - \frac{\omega_{cj}}{k_{y}L_{j}\omega} \right) \right]}{k^{2} - k^{2} V_{Fj}^{\prime 2} \left[\frac{k_{z}^{2}}{\omega^{2}} + \frac{k_{y}^{2}}{\omega^{2} - \omega_{cj}^{2}} \left(1 - \frac{\omega_{cj}}{k_{y}L_{j}\omega} \right) \right]},$$
 (10)

where $\omega_{pj} = (4\pi n_{j0}q_j^2/m_j)^{1/2}$ and $\omega_{cj} = q_j B_0/m_j c$ are the plasma frequency and the cyclotron frequency of the *j*th species, respectively. In (5), the scale length of inhomogeneity, $L_j = -n_{j0}/n'_{j0}$ with $n'_{j0} = \partial n_{j0}(x)/\partial x$. For quantum effect involving electron dynamics only, one can write $k^2 \lambda'_{Fe} = k^2 \lambda_{Fe}^2 + \omega_{Qe}^2/4\omega_{Pe}^2$, where $\lambda_{Fe} = V_{Fe}/\omega_{Pe}$ and $\omega_{Qe} = \hbar k^2/m_e$. We use (10) to find the general dielectric response function $\epsilon(\omega, \mathbf{k})$ of non-uniform quantum dusty magnetized plasma under various possible conditions,

$$\epsilon(\omega, \mathbf{k}) = 1 + \chi_e(\omega, \mathbf{k}) + \chi_i(\omega, \mathbf{k}) + \chi_d(\omega, \mathbf{k}).$$
(11)

It may be mentioned here that in the presence of considerable amount of neutral atoms/molecules in the dusty plasma, collision frequencies of plasma species with neutrals cannot be ignored. Also, the diamagnetic drift velocities of magnetized species in the non-uniform plasma may be important that might cause diamagnetic drift wave instabilities. In the presence of the ion diamagnetic drift velocity, $\mathbf{u}_{Di} = (c T_{Fi}/n_{i0}eB_0)(\partial n_{i0}(x)/\partial x)\hat{\mathbf{y}}$, and the average collision frequency, v_i , the ω in (11) should be properly taken into account by replacing ω with $\omega - k_y u_{Di} + iv_i$. We ignore diamagnetic drifts of electrons and dust grains because they are not assumed magnetized.

In order to study the diamagnetic DA and DLH drift wave instabilities in a collisional non-uniform quantum dusty plasma in the presence of an external static and homogeneous magnetic field, we assume

$$\omega \leqslant \omega_i^* \ll \omega_{ci}, \quad k_z^2 \ll k_y^2,$$
$$kV'_{Fi} \leqslant \omega \ll kV'_{Fe}, \tag{12}$$

where $\omega_i^* = \omega_{pi}^2 / k_y L_i \omega_{ci}$ is the drift frequency of cold and magnetized ions.

Dust-acoustic and DLH drift wave instabilities are presented in the following sections.

1. Dust-acoustic drift wave instabilities

$$(1/k^2 \lambda_{Fe}^{\prime 2} \simeq 1/k^2 \lambda_{Fe}^2 \gg f_i \gg 1)$$

First, we consider a high-density quantum dusty plasma where quantum effect arises through the Fermi degenerate pressure of electrons and the small Bohm potential is neglected. Here we consider that the electrons fully degenerate with the Fermi temperature T_{Fe} , ions are magnetized and collisional, and the dust grains are unmagnetized but collisional (Baines et al. 1965; Rosenberg and Shukla 2004). In this case, the dielectric function satisfying (12) is given by

$$\epsilon(\omega, \mathbf{k}) = 1 + \frac{1}{k^2 \lambda_{Fe}^2} + f_i - \frac{\omega_i^*}{\Omega_i + iv_i} - \frac{\omega_{pd}^2}{\omega(\omega + iv_d)}, \quad (13)$$

where $\Omega_i = \omega - k_y u_{Di}$ and $f_i = \omega_{pi}^2 / \omega_{ci}^2$; here u_{Di} is the magnetized ion fluid drift velocity. For $iv_i \ge \Omega_i$, $iv_d \le \omega$, and $1/k^2 \lambda_{Fe}^2 \ge f_i$, the above equation for $\epsilon = 0$ can be

written as

$$1 + \frac{i\omega_{pi}^2}{k_y L_i \omega_{ci}} \frac{k^2 \lambda_{Fe}^2}{v_i} - \frac{k^2 C_{Fd}^2}{\omega^2} \left(1 - \frac{iv_d}{\omega}\right) = 0, \quad (14)$$

where $C_{Fd}^2 = \omega_{pd}^2 V_{Fe}^2 / \omega_{pe}^2$. The relation $\epsilon_r = 0$, where ϵ_r is the real part of the dielectric function, yields the dispersion relation of the modified DA wave in the Fermi quantum dusty plasma

$$\omega_{DA} = k C_{Fd}.$$
 (15)

Assuming $k_y L_i \omega_{ci} < k C_{Fi}$, where $C_{Fi} = \omega_{pi} \lambda_{Fe}$ is the ionacoustic speed at the electron-Fermi temperature, (14) describes the DA drift instability given by

$$\frac{\omega}{\omega_{DA}} = \frac{1+i}{\sqrt{2}} \left(\frac{v_i |L_i| \omega_{ci}}{k_y C_{Fi}^2}\right)^{1/2} - \frac{i v_d}{2 \omega_{DA}},\qquad(16)$$

where $L_i = -|L_i|$ for a positive ion density gradient. We note that DA wave in the Fermi quantum dusty plasma grows in amplitude when the ion inhomogeneity scale length is sufficiently small.

Next, we consider the strong quantum effect arising through the Bohm potential compared with the effect arising through the Fermi degenerate pressure. This situation arises for a relatively low-density quantum plasma and short wavelength DA waves. In this case, the dispersion relation is obtained from

$$1 + \frac{4\omega_{pe}^2}{\omega_{Qe}^2} + f_i - \frac{\omega_i^*}{\Omega_i} \left(1 - \frac{iv_i}{\Omega_i}\right) - \frac{\omega_{pd}^2}{\omega^2} \left(1 - \frac{iv_d}{\omega}\right) = 0, (17)$$

where $\Omega_i > iv_i$ is made.

However, for $iv_i \ge \Omega_i$ and $iv_d \ll \omega$, (14) reduces to

$$1 - ix - \frac{\omega_{pd}^2 / \left(4\omega_{pe}^2 / \omega_{Qe}^2\right)}{\omega^2} \left(1 - \frac{iv_d}{\omega}\right) = 0, \qquad (18)$$

where

$$x = \frac{\delta\hbar^2 k^4}{4m_e m_i v_i} \frac{1}{k_y |L_i|\omega_{ci}},\tag{19}$$

with non-neutrality parameter in the dusty plasma, $\delta = n_{i0}/n_{e0}$. Assuming $\delta \hbar^2 k^4 \gg 4m_e m_i k_y |L_i| \omega_{ci} v_i$, the DA drift instability is described by

$$\frac{\omega}{\omega_{QDA}} = \frac{1+i}{\sqrt{2}} \left(\frac{4 m_e m_i v_i k_y |L_i| \omega_{ci}}{\delta \hbar^2 k^4} \right)^{1/2} - \frac{i v_d}{2 \omega_{QDA}}, \quad (20)$$

where the quantum DA wave frequency because of the Bohm potential is given by

$$\omega_{QDA} = \frac{\omega_{pd} \, \omega_{Qe}}{2 \, \omega_{pe}}$$
$$= \frac{\hbar k^2 Z_d}{2\sqrt{m_e m_d}} \sqrt{\frac{n_{d0}}{n_{e0}}} \,. \tag{21}$$

Here we note that the quantum modified DA wave suffers instability for small $v_i|L_i|$.

2. Dust-lower-hybrid drift wave instabilities $(f_i \ge 1/k^2 \lambda_{E_e}^2 \ge 1)$

For $iv_i \ge \Omega_i$ and $iv_d \ll \omega$ in this case, the dielectric function from (13) reduces to

$$\epsilon(\omega, \mathbf{k}) = 1 + i \left(\frac{\omega_{pi}^2}{k_y L_i v_i \omega_{ci} f_i} \right) - \frac{\omega_{dlh}^2}{\omega^2} \left(1 - \frac{i v_d}{\omega} \right), \quad (22)$$

where the DLH frequency is given by $\omega_{dlh} = \omega_{pd}\omega_{ci}/\omega_{pi}$. For a positive density gradient, $L_i = -|L_i|$, the dispersion relation for a growing DLH drift wave is obtained from (22) as,

$$1 - \frac{i\,\omega_{ci}}{v_i k_y |L_i|} - \frac{\omega_{dlh}^2}{\omega^2} \left(1 - \frac{iv_d}{\omega}\right) = 0.$$
(23)

Thus, the DLH instability for $\omega_{ci} \gg v_i k_v |L_i|$ is given by

$$\frac{\omega}{\omega_{dlh}} = \frac{1+i}{\sqrt{2}} \left(\frac{v_i k_y |L_i|}{\omega_{ci}}\right)^{1/2} - \frac{i v_d}{2\omega_{dlh}} \,. \tag{24}$$

Obviously, the long-wavelength DLH wave possesses instability in the inhomogeneous dusty plasma.

Next, we consider $iv_i \ll \Omega_i$ and $iv_d \ll \omega$. Thus, from (13) we have

$$1 + \frac{\omega_{ci}}{k_y |L_i|(\omega - k_y u_{Di})} - \frac{\omega_{dlh}^2}{\omega^2} = 0.$$
 (25)

For $k_y u_{Di} > \omega$, we obtain a purely growing DLH instability from

$$\omega^2 = -\frac{\omega_{dlh}^2 k_y^2 |L_i| u_{Di}}{\omega_{ci}}.$$
(26)

The growth rate of the purely growing DLH instability depends on the diamagnetic drift velocity of ions.

In summary, we have investigated the DA and DLH diamagnetic drift wave instabilities in the presence of a static ambient or an applied magnetic field in a collisional non-uniform quantum dusty plasma. For the high-density plasma, the quantum effect on DA drift instability arises through the Fermi degenerate pressure. However, for a relatively low-density quantum plasma (e.g., nano-scale microelectronics, semiconductor plasmas, or laser-produced plasmas) and for short wavelength, the quantum effect is taken through the Bohm potential of quantum plasma. Various possible instability conditions are found for the diamagnetic DA wave instabilities. The linear dispersion relations of DA wave are also found in quantum plasma (cf. Opher et al. 2001; Shukla and Stenflo 2006a). In the long wavelength approximation, we find conditions for the DLH diamagnetic ion drift instabilities where the quantum effect is not significant for the inhomogeneous quantum dusty plasma.

We would like to stress here that in quantum dusty magnetoplasmas, such as dense astrophysical environments, microelectronics, nano-structured materials, high-density laser-produced plasmas, etc., the spatial non-uniformity of parameters, particularly the plasma densities of species, may be obviously quite common. Equilibrium density non-uniformity in the presence of magnetic fields gives rise to electrostatic drift waves in quantum systems. Various possible instability conditions are found for electrostatic diamagnetic drift instabilities. These instabilities would excite possible electrostatic drift waves, which give rise to the fundamental properties of non-uniform quantum plasmas. Electromagnetic drift waves must also be explored in quantum dusty magnetoplasmas which are beyond the scope of the present paper. The nonlinear wave–wave interactions at large-amplitude appear to be of interest also giving new properties of non-uniform dusty quantum magnetoplasmas. The parametric cascading of electromagnetic drift waves at large amplitudes in the presence of electrostatic drift waves may also lead to the turbulence of quantum plasmas, giving rise to various nonlinear effects.

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