

SITELLE at the CFHT

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Abstract. SITELLE is the new imaging Fourier transform spectrograph of the Canada-France-Hawaii Telescope. It produces an impressive 4 million spectra in a single datacube in selected bandpasses from 350 to 900 nm. Its large FOV (11'x11') and its high spatial sampling (0.32"/pixel, seeing limited) allow us to study extended objects with an unprecedented view (Drissen *et al.* 2014). SITELLE's first observations of nearby galaxies revealed its capabilities to conduct detailed studies of emission line regions.

Keywords. Instrum.: Imaging Spectrograph; ISM Structure; HII Regions; Abundances

1. Summary

Three datacubes in the spectral bands SN1 [365 - 384 nm], SN2 [484 - 512 nm], and SN3 [648 - 685 nm] were collected on the well known face-on spiral NGC 628 during SITELLE's science verification run. Data reduction was performed using ORBS (Martin 2015). More than 3500 HII regions were detected in the disk. Figure 1 shows the reconstructed spectrum of one region. The wavelength and flux calibrations are well confirmed from comparison with data from the literature.

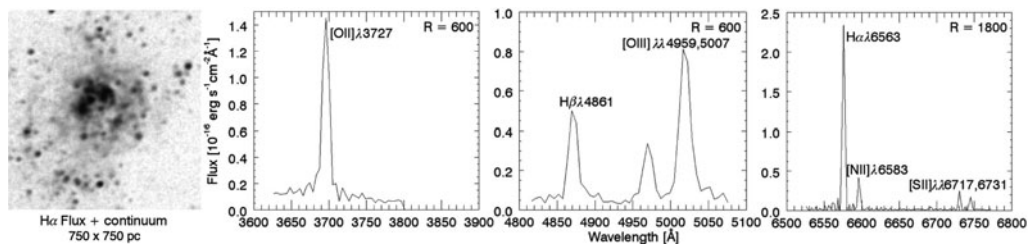


Figure 1. Details of an HII region in NGC 628, with its integrated spectrum from SITELLE.

With SITELLE, ionization structures in NGC 628 are resolved (~ 25 pc) and therefore, new diagnostic tools are required to better constrain the physical parameters (age, mass, abundances, ionization parameter, escaping photon fraction) of each star-forming region. The photoionization codes CLOUDY and MAPPINGS will be used to prepare a large database tailored for SITELLE in order to facilitate the comparison with different emission line ratios. A morphological classification is performed using constraints on the H α profiles. The additional information provided by the analysis of resolved ionization morphologies and kinematics will bring new clues on the content and the physical conditions of a very large population of HII regions, along with diffuse ionized gas in disk galaxies.

References

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